

CONSERVATION LAWS ON THE SPHERE: FROM SHALLOW WATER TO BURGERS

Matania Ben-Artzi

Institute of Mathematics, Hebrew University, Jerusalem, Israel

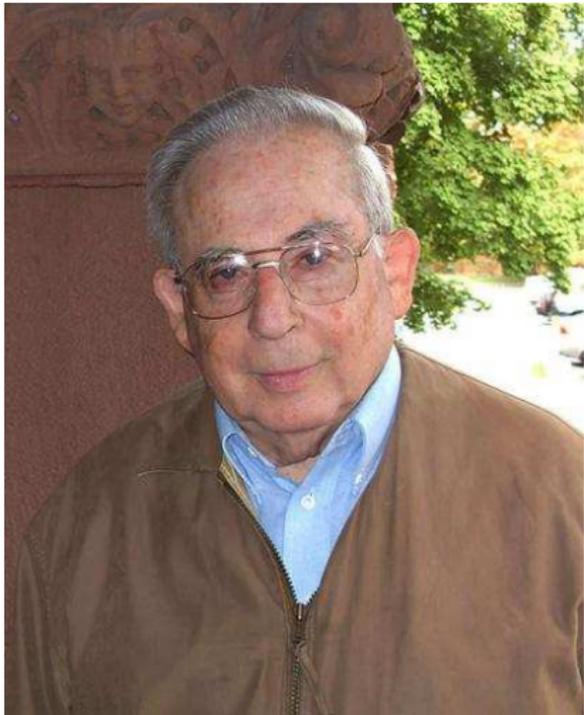
Advances in Applied Mathematics

IN MEMORIAM OF PROFESSOR SAUL ABARBANEL

Tel Aviv University

December 2018

joint work with JOSEPH FALCOVITZ, PHILIPPE LEFLOCH



“...together with David Gottlieb we noticed that some of the stuff that people were doing, the formulation was not strongly well posed, which is a mathematical point of view. So we got interested in how to make it more posed.” (Interview with P. Davis, Brown University, 2003).
“Problems should be studied in a ‘physico-mathematical’ fashion” –(private communication)

General Circulation Model –JETSTREAM

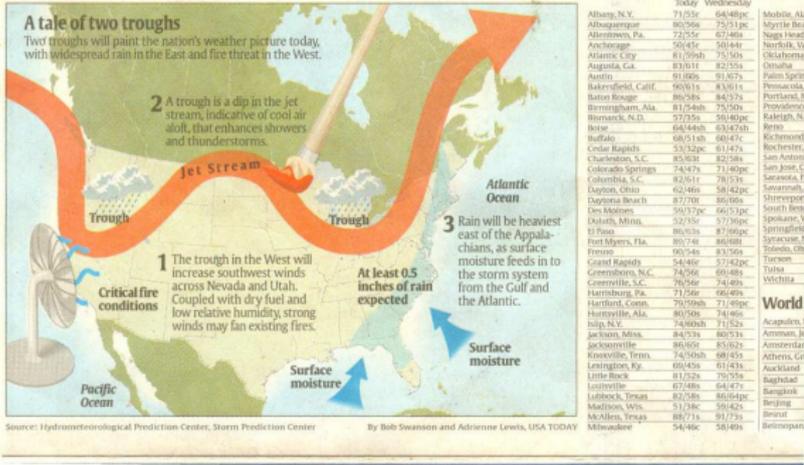
USA TODAY weather focus

Check out The Weather Guys, at blogs.usatoday.com/weather

National forecast

A tale of two troughs

Two troughs will paint the nation's weather picture today, with widespread rain in the East and fire threat in the West.



Source: Hydrometeorological Prediction Center, Storm Prediction Center

By Bob Swanson and Adrienne Lewis, USA TODAY

Today Wednesday

Albany, N.Y.	71/55F	64/48F
Albuquerque	80/25F	75/51F
Allentown, Pa.	72/25F	67/46F
Anchorage	50/47F	50/44F
Atlanta City	81/29F	75/50F
Augusta, Ga.	83/31F	82/55F
Austin	91/66F	91/67F
Bakersfield, Calif.	59/81F	83/81F
Baton Rouge	89/58F	84/57F
Birmingham, Ala.	81/54F	75/50F
Birmingham, N.D.	57/25F	50/40F
Boise	68/44F	63/47F
Buffalo	68/51F	60/47F
Cedar Rapids	53/32F	61/47F
Charleston, S.C.	85/63F	82/58F
Colorado Springs	74/47F	71/40F
Columbia, S.C.	82/51F	78/51F
Dayton, Ohio	62/48F	58/42F
Daytona Beach	81/70F	80/66F
Des Moines	59/37F	66/53F
Duluth, Minn.	32/25F	57/36F
El Paso	86/63F	87/66F
Fort Myers, Fla.	89/74F	86/68F
Fresno	88/64F	83/56F
Grand Rapids	54/40F	57/42F
Greensboro, N.C.	74/56F	69/48F
Greenville, S.C.	78/56F	74/49F
Harrisburg, Pa.	71/56F	66/49F
Hartford, Conn.	70/59F	71/49F
Huntsville, Ala.	80/59F	74/46F
Islip, N.Y.	74/60F	71/52F
Jackson, Miss.	84/53F	80/53F
Jacksonville	86/60F	85/62F
Knoxville, Tenn.	74/50F	68/45F
Lexington, Ky.	68/45F	61/43F
Little Rock	81/52F	75/55F
Louisville	67/48F	64/47F
Lubbock, Texas	67/58F	66/54F
Madison, Wis.	51/38F	59/42F
Madison, Texas	88/71F	91/73F
Madison, Wis.	54/40F	58/40F

Mobile, Ala.	
Myrtle Beach	
Nags Head	
Norfolk, Va.	
Ocala/Orlando	
Omaha	
Palm Springs	
Pensacola, Fla.	
Portland, Me.	
Providence	
Raleigh, N.C.	
RENO	
Richmond	
Rochester, N.Y.	
San Antonio	
San Jose, Ca.	
Sarasota, Fl.	
Savannah, Ga.	
Shreveport	
South Bend	
Spokane, Wa.	
Springfield	
Syracuse, N.Y.	
Tulsa, Okla.	
Tucson	
Tulsa	
Wichita	

World

Acapulco, J.	
Amman, Jo.	
Amsterdam	
Athens, Gr.	
Auckland	
Bahia, Br.	
Bangkok	
Beijing	
Bombay	
Buenos Aires	
Birmingham	

General Circulation Model - JETSTREAM

USA TODAY weather focus

Check out The Weather Guys, at blogs.usatoday.com/weather

National forecast

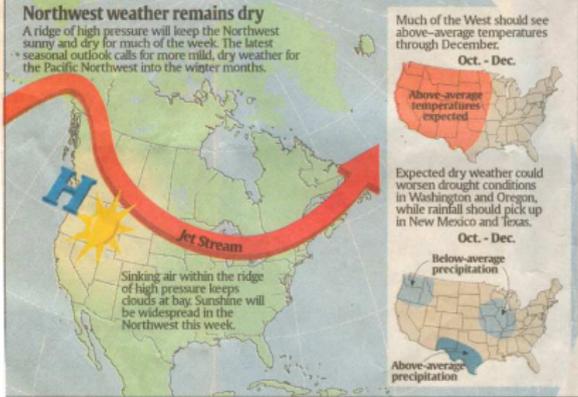


Today	Wednesday
Albany, N.Y.	71/55F 64/49F
Albuquerque	80/25F 72/21F
Allentown, Pa.	72/25F 67/46F
Anchorage	50/44F 50/44F
Atlanta City	81/29F 75/50F
Augusta, Ga.	83/61F 82/55F
Austin	91/66F 91/67F
Bakersfield, Calif.	90/61F 83/61F
Baton Rouge	89/58F 84/37F
Birmingham, Ala.	81/24F 75/20F
Bismarck, N.D.	57/25F 50/40F
Boise	68/44F 63/47F
Buffalo	68/51F 60/47F
Cedar Rapids	53/32F 61/47F
Charleston, S.C.	85/63F 82/38F
Colorado Springs	74/47F 71/40F
Columbia, S.C.	82/51F 78/51F
Dayton, Ohio	62/46F 58/42F
Daytona Beach	81/70F 80/66F
Des Moines	59/37F 66/53F
Duluth, Minn.	52/25F 57/36F
El Paso	86/65F 87/66F
Fort Myers, Fla.	89/74F 86/68F
Fort Worth	80/54F 83/50F
Grand Rapids	54/40F 57/42F
Greensboro, N.C.	74/56F 69/48F
Greenville, S.C.	78/56F 74/49F
Harrisburg, Pa.	71/56F 66/40F
Hartford, Conn.	70/59F 71/49F
Huntsville, Ala.	80/59F 74/46F
Islip, N.Y.	74/60F 71/52F
Jackson, Miss.	84/53F 80/53F
Jacksonville	86/60F 85/62F
Knoxville, Tenn.	74/50F 68/45F
Lexington, Ky.	68/45F 61/43F
Little Rock	81/52F 75/55F
Louisville	67/48F 64/47F
Lubbock, Texas	67/50F 60/54F
Madison, Wis.	51/38F 59/42F
Madison, Vt.	88/71F 91/73F
Manassas	54/46F 58/40F
Mobile, Ala.	
Myrtle Beach	
Napa, Calif.	
Norfolk, Va.	
Ocala, Fla.	
Omaha	
Palm Spring	
Pensacola, Fla.	
Portland, Me.	
Providence	
Raleigh, N.C.	
RENO	
Richmond	
Rochester	
San Antonio	
San Jose, Calif.	
Sarasota, Fla.	
Savannah, Ga.	
Shreveport	
Springfield	
Syracuse, N.Y.	
Tulsa, Okla.	
Victoria	
Wichita	
World	
Acapulco	
Amman, Jord.	
Amsterdam	
Athens, Greece	
Auckland	
Bangkok	
Beijing	
Bombay	
Buenos Aires	
Birmingham	

USA TODAY weather focus

Check out The Weather Guys, at blogs.usatoday.com/weather

Natio



Albany, N.Y.	
Albuquerque	
Allentown, Pa.	
Anchorage	
Atlanta City	
Augusta, Ga.	
Austin	
Bakersfield	
Baton Rouge	
Birmingham	
Bismarck	
Boise	
Buffalo	
Cedar Rapids	
Charleston	
Colorado Springs	
Columbia	
Dayton, Ohio	
Daytona Beach	
Des Moines	
Duluth, Minn.	
El Paso	
Fort Myers	
Fort Worth	
Grand Rapids	
Greensboro	
Greenville	
Harrisburg	
Hartford	
Huntsville	
Islip, N.Y.	
Jackson, Miss.	
Jacksonville	
Knoxville	
Lexington	
Little Rock	
Louisville	
Lubbock, Texas	
Madison, Wis.	
Madison, Vt.	
Manassas	

source: Climate Prediction Center

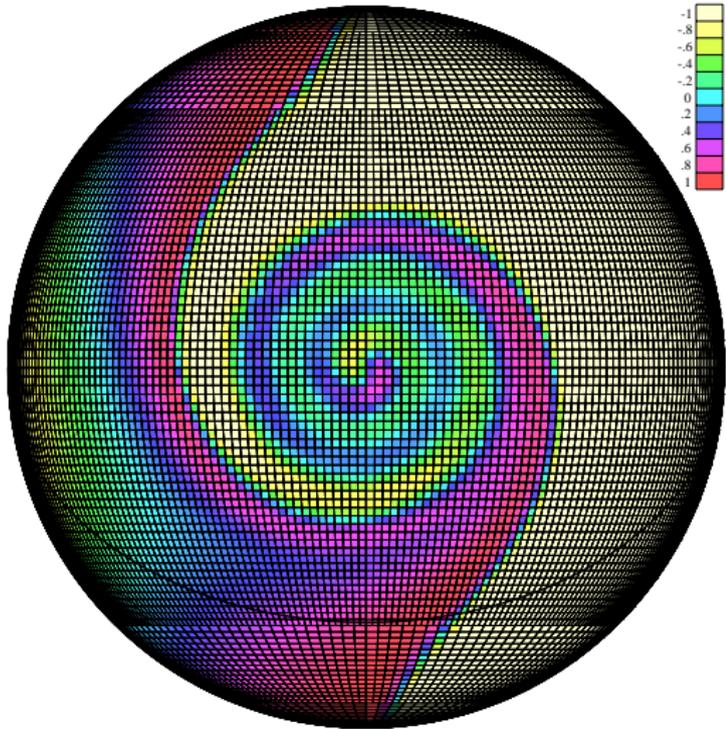
By Bob Swanson and Adrienne Lewis, USA TODAY

SEPTEMBER 26, 2006

MOVING VORTEX

R.D. Nair and C. Jablonowski—Moving vortices on the sphere: A test case for horizontal advection problems,

Monthly Weather Review 136(2008)699–711

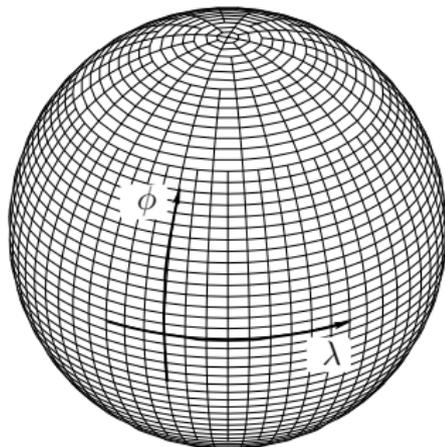


SPHERE OF f - Oct. Mon, Apr. 13, 2003 12:20:16

Grid on sphere—the Kurihara Grid

D. J. Williamson—The evolution of dynamical cores for global atmospheric models,
Journal of the meteorological society of Japan 85B (2007)241–269

DISCUSSION: The “POLE PROBLEM”



SOME REFERENCES–GEOPHYSICAL

P.S. Marcus–Numerical simulation of Jupiter's great red spot ,
Nature (1988)

SOME REFERENCES–GEOPHYSICAL

P.S. Marcus–Numerical simulation of Jupiter's great red spot ,
Nature (1988)

J.Y-K. Cho, and L. M. Polvani–The morphogenesis of bands and
zonal winds in the atmospheres on the giant outer planets ,
Science (1996)

SOME REFERENCES–GEOPHYSICAL

P.S. Marcus–Numerical simulation of Jupiter's great red spot ,
Nature (1988)

J.Y-K. Cho, and L. M. Polvani–The morphogenesis of bands and
zonal winds in the atmospheres on the giant outer planets ,
Science (1996)

J. Galewski, R.K. Scott and L. M. Polvani–An initial-value problem
for testing numerical models of the global shallow water equations
, Tellus (2004)

SOME REFERENCES–GEOPHYSICAL

P.S. Marcus–Numerical simulation of Jupiter's great red spot ,
Nature (1988)

J.Y-K. Cho, and L. M. Polvani–The morphogenesis of bands and
zonal winds in the atmospheres on the giant outer planets ,
Science (1996)

J. Galewski, R.K. Scott and L. M. Polvani–An initial-value problem
for testing numerical models of the global shallow water equations
, Tellus (2004)

T. Woollings and M. Blackburn–The North Atlantic jet stream
under climate change and its relation to the NAO and EA patterns
, (NAO=North Atlantic Oscillations,EA=East Atlantic)
Journal of Climate (2012)

SOME REFERENCES–COMPUTATIONAL

D. L. Williamson, J.B. Drake, J.J. Hack, R. Jakob and P. N. Swarztrauber—A standard test set for numerical approximations to the shallow water equations in spherical geometry, J. Comp. Physics (1992)

SOME REFERENCES–COMPUTATIONAL

D. L. Williamson, J.B. Drake, J.J. Hack, R. Jakob and P. N. Swarztrauber—A standard test set for numerical approximations to the shallow water equations in spherical geometry, J. Comp. Physics (1992)

J. A. Rossmannith, D. S. Bale and R. J. LeVeque—A wave propagation method for hyperbolic systems on curved manifolds , J. Comp. Physics (2004)

SOME REFERENCES–COMPUTATIONAL

D. L. Williamson, J.B. Drake, J.J. Hack, R. Jakob and P. N. Swarztrauber—A standard test set for numerical approximations to the shallow water equations in spherical geometry, J. Comp. Physics (1992)

J. A. Rossmannith, D. S. Bale and R. J. LeVeque—A wave propagation method for hyperbolic systems on curved manifolds , J. Comp. Physics (2004)

P. A. Ullrich, C. Jablonowski and B. van Leer—High-order finite-volume methods for the shallow water equations on the sphere , J. Comp. Physics (2010)

SOME REFERENCES–COMPUTATIONAL

D. L. Williamson, J.B. Drake, J.J. Hack, R. Jakob and P. N. Swarztrauber—A standard test set for numerical approximations to the shallow water equations in spherical geometry, J. Comp. Physics (1992)

J. A. Rossmanith, D. S. Bale and R. J. LeVeque—A wave propagation method for hyperbolic systems on curved manifolds , J. Comp. Physics (2004)

P. A. Ullrich, C. Jablonowski and B. van Leer—High-order finite-volume methods for the shallow water equations on the sphere , J. Comp. Physics (2010)

L. Bao, R.D. Nair and H.M. Tufu—A mass and momentum flux-form high-order discontinuous Galerkin shallow water model on the cubed-sphere , J. Comp. Physics (2013)

SOME BOOKS

D. R. Durran—Numerical Methods for Fluid Dynamics: With Applications to Geophysics , Springer (1999,2010)

N. Paldor—Shallow Water Waves on the Rotating Earth , Springer (2015)

R. Salmon—Lectures on Geophysical Fluid Dynamics, Oxford University Press (1998)

GRP METHODOLOGY USING RIEMANN INVARIANTS and GEOMETRIC COMPATIBILITY

J. Li and G. Chen—The generalized Riemann problem method for the shallow water equations with topography, Int. J. Numer. Methods in Engineering(2006)

M. Ben-Artzi, J. Li and G. Warnecke—A direct Eulerian GRP scheme for compressible fluid flows, J. Comp. Physics (2006)

M. Ben-Artzi, J. Falcovitz and Ph. LeFloch— Hyperbolic conservation laws on the sphere: A geometry compatible finite volume scheme, J.Comp. Physics (2009)

DERIVATION OF THE MODEL

INVARIANT FORM

TWO SYSTEMS

Lower-case letters = Inertial system

Capital letters = Rotating system .

Time derivatives of vector functions: $\dot{\vec{q}}(t), \ddot{\vec{Q}}(t)$.

Connection by **ROTATION MATRIX**

$$\vec{x} = R(t)\vec{X}.$$

$$\dot{\vec{x}} = \frac{d}{dt}R(t)\vec{X} = R(t)(\vec{\Omega} \times \vec{X}).$$

$\vec{\Omega} = \vec{\Omega}(t)$ = angular velocity in the rotating system. It is *constant* (namely, independent of time) in the rotating system.

If $\vec{X}(t)$ represents a moving particle in the rotating system,

$$\dot{\vec{x}} = \frac{d}{dt}R(t)\vec{X} = R(t)(\vec{\Omega} \times \vec{X}) + R(t)(\dot{\vec{X}}).$$

If $\vec{X}(t)$ represents a moving particle in the rotating system,

$$\dot{\vec{x}} = \frac{d}{dt}R(t)\vec{X} = R(t)(\vec{\Omega} \times \vec{X}) + R(t)(\dot{\vec{X}}).$$

$$\begin{aligned}\ddot{\vec{x}} &= R(t)\{\vec{\Omega} \times (\vec{\Omega} \times \vec{X}) + \vec{\Omega} \times \dot{\vec{X}} + \vec{\Omega} \times \dot{\vec{X}} + \ddot{\vec{X}}\} \\ &= R(t)\{\vec{\Omega} \times (\vec{\Omega} \times \vec{X}) + 2\vec{\Omega} \times \dot{\vec{X}} + \ddot{\vec{X}}\}.\end{aligned}$$

If $\vec{X}(t)$ represents a moving particle in the rotating system,

$$\dot{\vec{x}} = \frac{d}{dt}R(t)\vec{X} = R(t)(\vec{\Omega} \times \vec{X}) + R(t)(\dot{\vec{X}}).$$

$$\begin{aligned}\ddot{\vec{x}} &= R(t)\{\vec{\Omega} \times (\vec{\Omega} \times \vec{X}) + \vec{\Omega} \times \dot{\vec{X}} + \vec{\Omega} \times \dot{\vec{X}} + \ddot{\vec{X}}\} \\ &= R(t)\{\vec{\Omega} \times (\vec{\Omega} \times \vec{X}) + 2\vec{\Omega} \times \dot{\vec{X}} + \ddot{\vec{X}}\}.\end{aligned}$$

Particle of mass m , force \vec{f} (in the inertial system):

$$R(t)(m\ddot{\vec{X}}) = \vec{f} - mR(t)\{\vec{\Omega} \times (\vec{\Omega} \times \vec{X}) + 2\vec{\Omega} \times \dot{\vec{X}}\}.$$

Lagrangian formulation: particle has *unit mass*, and is an element of a fluid continuum moving (approximately) on the spherical surface of the earth S .

\vec{N} = outward unit normal on the sphere S .

TWO “ADDITIONAL FORCES”

CENTRIFUGAL FORCE $\vec{\Omega} \times (\vec{\Omega} \times \vec{X})$

CORIOLIS FORCE $2\vec{\Omega} \times \dot{\vec{X}}$

Velocity $\vec{V} = \dot{\vec{X}}$.

ASSUMPTION I:

There are two body forces acting on the particle:

- ▶ $-\vec{G}$ — the gravity force .
- ▶ \vec{H} — the hydrostatic force (due to fluid pressure).

Total force (in the inertial system) on the unit mass is

$$\vec{f} = R(t)(-\vec{G} + \vec{H}).$$

ASSUMPTION I:

There are two body forces acting on the particle:

- ▶ $-\vec{G}$ — the gravity force .
- ▶ \vec{H} — the hydrostatic force (due to fluid pressure).

Total force (in the inertial system) on the unit mass is

$$\vec{f} = R(t)(-\vec{G} + \vec{H}).$$

$$R(t)(\ddot{\vec{X}}) = R(t)\{-\vec{G} + \vec{H} - \vec{\Omega} \times (\vec{\Omega} \times \vec{X}) - 2\vec{\Omega} \times \dot{\vec{X}}\}.$$

ASSUMPTION I:

There are two body forces acting on the particle:

- ▶ $-\vec{G}$ — the gravity force .
- ▶ \vec{H} — the hydrostatic force (due to fluid pressure).

Total force (in the inertial system) on the unit mass is

$$\vec{f} = R(t)(-\vec{G} + \vec{H}).$$

$$R(t)(\ddot{\vec{X}}) = R(t)\{-\vec{G} + \vec{H} - \vec{\Omega} \times (\vec{\Omega} \times \vec{X}) - 2\vec{\Omega} \times \dot{\vec{X}}\}.$$

Note: \vec{X} is a three-dimensional vector in the rotational system.

Later: Confine to the sphere $S : r = a$, by assuming that the fluid volume is very “thin” (vertically).

ASSUMPTION I:

There are two body forces acting on the particle:

- ▶ $-\vec{G}$ — the gravity force .
- ▶ \vec{H} — the hydrostatic force (due to fluid pressure).

Total force (in the inertial system) on the unit mass is

$$\vec{f} = R(t)(-\vec{G} + \vec{H}).$$

$$R(t)(\ddot{\vec{X}}) = R(t)\{-\vec{G} + \vec{H} - \vec{\Omega} \times (\vec{\Omega} \times \vec{X}) - 2\vec{\Omega} \times \dot{\vec{X}}\}.$$

Note: \vec{X} is a three-dimensional vector in the rotational system.

Later: Confine to the sphere $S : r = a$, by assuming that the fluid volume is very “thin” (vertically).

$$\ddot{\vec{X}} = -\vec{G} + \vec{H} - \vec{\Omega} \times (\vec{\Omega} \times \vec{X}) - 2\vec{\Omega} \times \dot{\vec{X}}.$$

ASSUMPTION II:

Some constant $g^* > 0$,

$$-\vec{G} - \vec{\Omega} \times (\vec{\Omega} \times \vec{X}) = -g^* \vec{N}.$$

ASSUMPTION II:

Some constant $g^* > 0$,

$$-\vec{G} - \vec{\Omega} \times (\vec{\Omega} \times \vec{X}) = -g^* \vec{N}.$$

MEANING: Earth is not a perfect sphere, the combination of the gravitational and the centrifugal forces can be incorporated into a perfect spherical setting where the “modified” gravitational force is radial.

GEOPHYSICAL LITERATURE: [geopotential](#) and the [geoid](#).

ASSUMPTION II:

Some constant $g^* > 0$,

$$-\vec{G} - \vec{\Omega} \times (\vec{\Omega} \times \vec{X}) = -g^* \vec{N}.$$

MEANING: Earth is not a perfect sphere, the combination of the gravitational and the centrifugal forces can be incorporated into a perfect spherical setting where the “modified” gravitational force is radial.

GEOPHYSICAL LITERATURE: [geopotential](#) and the [geoid](#).

Remain (apart from the modified gravitation):

CORIOLIS FORCE $-2\vec{\Omega} \times \vec{V}$ HYDROSTATIC FORCE \vec{H} .

ASSUMPTION II:

Some constant $g^* > 0$,

$$-\vec{G} - \vec{\Omega} \times (\vec{\Omega} \times \vec{X}) = -g^* \vec{N}.$$

MEANING: Earth is not a perfect sphere, the combination of the gravitational and the centrifugal forces can be incorporated into a perfect spherical setting where the “modified” gravitational force is radial.

GEOPHYSICAL LITERATURE: [geopotential](#) and the [geoid](#).

Remain (apart from the modified gravitation):

CORIOLIS FORCE $-2\vec{\Omega} \times \vec{V}$ HYDROSTATIC FORCE \vec{H} .

Earth surface $S : r = a$: At every point orthonormal system (fixed in rotational system): unit normal \vec{N} + “tangential plane”.

$$\vec{X} = X_N \vec{N} + \vec{X}_T$$

ASSUMPTION II:

Some constant $g^* > 0$,

$$-\vec{G} - \vec{\Omega} \times (\vec{\Omega} \times \vec{X}) = -g^* \vec{N}.$$

MEANING: Earth is not a perfect sphere, the combination of the gravitational and the centrifugal forces can be incorporated into a perfect spherical setting where the “modified” gravitational force is radial.

GEOPHYSICAL LITERATURE: [geopotential](#) and the [geoid](#).

Remain (apart from the modified gravitation):

CORIOLIS FORCE $-2\vec{\Omega} \times \vec{V}$ HYDROSTATIC FORCE \vec{H} .

Earth surface $S : r = a$: At every point orthonormal system (fixed in rotational system): unit normal \vec{N} + “tangential plane”.

$$\vec{X} = X_N \vec{N} + \vec{X}_T$$

$$(\ddot{\vec{X}})_T = \vec{H}_T - (2\vec{\Omega} \times \vec{V})_T.$$

SHALLOW-WATER MODEL

Incompressible fluid occupies a “thin”, yet varying in depth (and in time) layer above the spherical surface $S : r = a$.

SHALLOW-WATER MODEL

Incompressible fluid occupies a “thin”, yet varying in depth (and in time) layer above the spherical surface $S : r = a$.

Y = point on the sphere,

z = vertical distance (along the normal \vec{N}) from the surface

$S : z = 0$.

$$0 \leq z \leq h(Y, t), \quad Y \in S.$$

SHALLOW-WATER MODEL

Incompressible fluid occupies a “thin”, yet varying in depth (and in time) layer above the spherical surface $S : r = a$.

Y = point on the sphere,

z = vertical distance (along the normal \vec{N}) from the surface

$S : z = 0$.

$$0 \leq z \leq h(Y, t), \quad Y \in S.$$

“free surface” $z = h(Y, t)$ one of unknowns in the model .

SHALLOW-WATER MODEL

Incompressible fluid occupies a “thin”, yet varying in depth (and in time) layer above the spherical surface $S : r = a$.

Y = point on the sphere,

z = vertical distance (along the normal \vec{N}) from the surface

$S : z = 0$.

$$0 \leq z \leq h(Y, t), \quad Y \in S.$$

“free surface” $z = h(Y, t)$ **one of unknowns in the model** .

Fluid is incompressible (of unit density).

$h(Y, t)$ = height of column over $Y =$ (surface) mass density at Y, t .

ASSUMPTION III:

Tangential velocity $\vec{V}_T = \dot{\vec{X}}_T$ independent of the height z .

ASSUMPTION III:

Tangential velocity $\vec{V}_T = \dot{\vec{X}}_T$ independent of the height z .

Motion of “surface mass”, density h , determined by \vec{V}_T .

Conservation of mass:

$$\frac{\partial h}{\partial t}(Y, t) + \nabla_T \cdot (h(Y, t)\vec{V}_T) = 0.$$

ASSUMPTION III:

Tangential velocity $\vec{V}_T = \dot{\vec{X}}_T$ independent of the height z .

Motion of “surface mass”, density h , determined by \vec{V}_T .

Conservation of mass:

$$\frac{\partial h}{\partial t}(Y, t) + \nabla_T \cdot (h(Y, t)\vec{V}_T) = 0.$$

“Surface Lagrangian” derivative:

$$\frac{d}{dt} = \frac{\partial}{\partial t} + \vec{V}_T \cdot \nabla_T,$$

$$\frac{dh}{dt}(Y, t) = -h(Y, t)\nabla_T \cdot \vec{V}_T.$$

ASSUMPTION III:

Tangential velocity $\vec{V}_T = \dot{\vec{X}}_T$ independent of the height z .

Motion of “surface mass”, density h , determined by \vec{V}_T .

Conservation of mass:

$$\frac{\partial h}{\partial t}(Y, t) + \nabla_T \cdot (h(Y, t)\vec{V}_T) = 0.$$

“Surface Lagrangian” derivative:

$$\frac{d}{dt} = \frac{\partial}{\partial t} + \vec{V}_T \cdot \nabla_T,$$

$$\frac{dh}{dt}(Y, t) = -h(Y, t)\nabla_T \cdot \vec{V}_T.$$

Total derivative $\frac{d}{dt}$ is a “*surface derivative*”.

ASSUMPTION III:

Tangential velocity $\vec{V}_T = \dot{\vec{X}}_T$ independent of the height z .

Motion of “surface mass”, density h , determined by \vec{V}_T .

Conservation of mass:

$$\frac{\partial h}{\partial t}(Y, t) + \nabla_T \cdot (h(Y, t)\vec{V}_T) = 0.$$

“Surface Lagrangian” derivative:

$$\frac{d}{dt} = \frac{\partial}{\partial t} + \vec{V}_T \cdot \nabla_T,$$

$$\frac{dh}{dt}(Y, t) = -h(Y, t)\nabla_T \cdot \vec{V}_T.$$

Total derivative $\frac{d}{dt}$ is a “*surface derivative*”.

“Convective” part = $\vec{V}_T \cdot \nabla_T = \nabla_{V_T}$, **covariant derivative.**

ASSUMPTION IV:

Hydrostatic force \vec{H} is the gradient (in the rotating system) of the hydrostatic pressure in the fluid,

$$\vec{H} = -\nabla P.$$

ASSUMPTION IV:

Hydrostatic force \vec{H} is the gradient (in the rotating system) of the hydrostatic pressure in the fluid,

$$\vec{H} = -\nabla P.$$

$$P(Y, z = h(Y, t), t) = 0.$$

ASSUMPTION IV:

Hydrostatic force \vec{H} is the gradient (in the rotating system) of the hydrostatic pressure in the fluid,

$$\vec{H} = -\nabla P.$$

$$P(Y, z = h(Y, t), t) = 0.$$

Normal component of

$$\ddot{\vec{X}} = -\vec{G} + \vec{H} - \vec{\Omega} \times (\vec{\Omega} \times \vec{X}) - 2\vec{\Omega} \times \dot{\vec{X}}.$$

$$(\ddot{\vec{X}})_z = -g^* - \frac{\partial}{\partial z} P - (2\vec{\Omega} \times \vec{V})_z.$$

$$(\ddot{\vec{X}})_z = -g^* - \frac{\partial}{\partial z} P - (2\vec{\Omega} \times \vec{V})_z.$$

This equation is *three-dimensional*, $\{0 \leq z \leq h(Y, t), Y \in S\}$. In particular, the z - component of the Lagrangian derivative

$$(\dot{\vec{V}})_z \neq \frac{d}{dt} V_z, \quad V_z = \vec{V}_z \cdot \vec{N}.$$

$$(\ddot{\vec{X}})_z = -g^* - \frac{\partial}{\partial z} P - (2\vec{\Omega} \times \vec{V})_z.$$

This equation is *three-dimensional*, $\{0 \leq z \leq h(Y, t), Y \in S\}$. In particular, the z - component of the Lagrangian derivative

$$(\dot{\vec{V}})_z \neq \frac{d}{dt} V_z, \quad V_z = \vec{V}_z \cdot \vec{N}.$$

$$(2\vec{\Omega} \times \vec{V})_z = (2\vec{\Omega}_T \times \vec{V}_T)_z = (2\vec{\Omega}_T \times \vec{V}_T) \cdot \vec{N}.$$

$$(\ddot{\vec{X}})_z = -g^* - \frac{\partial}{\partial z} P - (2\vec{\Omega} \times \vec{V})_z.$$

This equation is *three-dimensional*, $\{0 \leq z \leq h(Y, t), Y \in S\}$. In particular, the z -component of the Lagrangian derivative

$$(\dot{\vec{V}})_z \neq \frac{d}{dt} V_z, \quad V_z = \vec{V}_z \cdot \vec{N}.$$

$$(2\vec{\Omega} \times \vec{V})_z = (2\vec{\Omega}_T \times \vec{V}_T)_z = (2\vec{\Omega}_T \times \vec{V}_T) \cdot \vec{N}.$$

Traditional treatment:

$$-g^* - \frac{\partial}{\partial z} P = 0, \quad \vec{\Omega}_T = 0.$$

$$V_z(Y, 0, t) \equiv 0 \Rightarrow \left[\dot{\vec{V}} \right]_z(Y, 0, t) \equiv 0.$$

R. Salmon writes: :

“In the traditional approximation, we neglect the horizontal component of the Earth’s rotation vector. This neglect has no convincing general justification; it must be justified in particular cases.”

We only use Assumption IV, in particular P is not assumed to vary linearly with respect to z .

$$V_z(Y, 0, t) \equiv 0 \Rightarrow \left[\dot{\vec{V}} \right]_z(Y, 0, t) \equiv 0.$$

$$V_z(Y, h(Y, t), t) = \frac{dh}{dt}.$$

$$(\dot{\vec{V}})_z = -g^* - \frac{\partial}{\partial z} P - (2\vec{\Omega} \times \vec{V})_z.$$

$$\int_0^{h(Y,t)} \left[\dot{\vec{V}} \right]_z dz = -[g^* + (2\vec{\Omega}_T \times \vec{V}_T)_z] h(Y,t) + P(Y,0,t),$$

$\vec{\Omega}_T$ depends only on Y , \vec{V}_T depends only on (Y,t)
(ASSUMPTION III).

$$P(Y,0,t) = [g^* + (2\vec{\Omega}_T \times \vec{V}_T)_z] h(Y,t), \quad Y \in S.$$

EFFECT of ROTATION on HYDROSTATIC PRESSURE!

$$P(Y, 0, t) = [g^* + (2\vec{\Omega}_T \times \vec{V}_T)_z]h(Y, t), \quad Y \in S.$$

BACK TO TANGENTIAL MOTION

$$(\ddot{\vec{X}})_T = (\dot{\vec{V}})_T = \vec{H}_T - (2\vec{\Omega} \times \vec{V})_T.$$

$$P(Y, 0, t) = [g^* + (2\vec{\Omega}_T \times \vec{V}_T)_z]h(Y, t), \quad Y \in S.$$

BACK TO TANGENTIAL MOTION

$$(\ddot{\vec{X}})_T = (\dot{\vec{V}})_T = \vec{H}_T - (2\vec{\Omega} \times \vec{V})_T.$$

$$\vec{H}_T = -\nabla_T P \Rightarrow$$

$$\frac{d}{dt} \vec{V}_T = -\nabla_T \left\{ [g^* + (2\vec{\Omega}_T \times \vec{V}_T)_z]h(Y, t) \right\} - (2\vec{\Omega} \times \vec{V})_T.$$

$$(2\vec{\Omega} \times \vec{V})_T = 2\vec{\Omega}_N \times \vec{V}_T.$$

$$P(Y, 0, t) = [g^* + (2\vec{\Omega}_T \times \vec{V}_T)_z]h(Y, t), \quad Y \in S.$$

BACK TO TANGENTIAL MOTION

$$(\ddot{\vec{X}})_T = (\dot{\vec{V}})_T = \vec{H}_T - (2\vec{\Omega} \times \vec{V})_T.$$

$$\vec{H}_T = -\nabla_T P \Rightarrow$$

$$\frac{d}{dt} \vec{V}_T = -\nabla_T \left\{ [g^* + (2\vec{\Omega}_T \times \vec{V}_T)_z]h(Y, t) \right\} - (2\vec{\Omega} \times \vec{V})_T.$$

$$(2\vec{\Omega} \times \vec{V})_T = 2\vec{\Omega}_N \times \vec{V}_T.$$

$$\frac{d}{dt} \vec{V}_T = -\nabla_T \left\{ [g^* + (2\vec{\Omega}_T \times \vec{V}_T)_z]h(Y, t) \right\} - 2\vec{\Omega}_N \times \vec{V}_T.$$

INVARIANT SHALLOW-WATER EQUATIONS ON THE SPHERE

$$\frac{dh}{dt}(Y, t) = -h(Y, t)\nabla_T \cdot \vec{V}_T.$$

$$\frac{d}{dt}\vec{V}_T = -\nabla_T \left\{ [g^* + (2\vec{\Omega}_T \times \vec{V}_T)_z]h(Y, t) \right\} - 2\vec{\Omega}_N \times \vec{V}_T.$$

INVARIANT SHALLOW-WATER EQUATIONS ON THE SPHERE

$$\frac{dh}{dt}(Y, t) = -h(Y, t) \nabla_T \cdot \vec{V}_T.$$

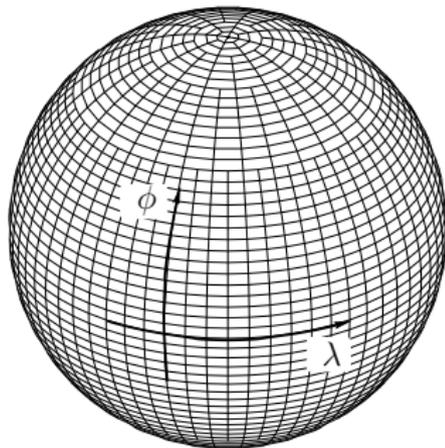
$$\frac{d}{dt} \vec{V}_T = -\nabla_T \left\{ [g^* + (2\vec{\Omega}_T \times \vec{V}_T)_z] h(Y, t) \right\} - 2\vec{\Omega}_N \times \vec{V}_T.$$

Compare Equator and Poles !

THE SW EQUATIONS – SPHERICAL COORDINATES

$$-\frac{\pi}{2} \leq \phi \leq \frac{\pi}{2}, \quad 0 \leq \lambda \leq 2\pi.$$

$$\frac{\partial h}{\partial t} + \frac{u}{a \cos \phi} \frac{\partial h}{\partial \lambda} + \frac{v}{a} \frac{\partial h}{\partial \phi} + \frac{h}{a \cos \phi} \left(\frac{\partial u}{\partial \lambda} + \cos \phi \frac{\partial v}{\partial \phi} \right) = \frac{hv \sin \phi}{a \cos \phi}.$$



$\delta = 0 \Rightarrow$ set $\vec{\Omega}_T = 0$, otherwise $\delta = 1$.

$$\begin{aligned} \frac{\partial u}{\partial t} + \frac{u - 2\delta\Omega h \cos \phi}{a \cos \phi} \frac{\partial u}{\partial \lambda} + \frac{v}{a} \frac{\partial u}{\partial \phi} + \frac{g^* - 2\delta\Omega u \cos \phi}{a \cos \phi} \frac{\partial h}{\partial \lambda} \\ = v \sin \phi \left\{ \frac{u}{a \cos \phi} + 2\Omega \right\} \quad , \end{aligned}$$

$$\begin{aligned} \frac{\partial v}{\partial t} + \frac{u}{a \cos \phi} \frac{\partial v}{\partial \lambda} + \frac{v}{a} \frac{\partial v}{\partial \phi} + \frac{g^* - 2\delta\Omega u \cos \phi}{a} \frac{\partial h}{\partial \phi} - \frac{2\delta\Omega h \cos \phi}{a} \frac{\partial u}{\partial \phi} \\ + 2\Omega \sin \phi \frac{\delta h u}{a} = -\frac{u^2}{a \cos \phi} \sin \phi - 2\Omega u \sin \phi \quad . \end{aligned}$$

THE SPLIT SCHEME WITH SOURCE TERMS

$$\psi_t = A[\psi] + B[\psi] + f(\cdot, \psi),$$

Consider first the homogeneous evolution

$$\psi_t = A[\psi] + B[\psi],$$

$$\psi(t) = \mathfrak{L}_{AB}(t)\psi_0.$$

Nonhomogeneous system: A, B are **linear**, but not necessarily commuting, the solution is expressed by the Duhamel principle

$$\psi(t) = \mathfrak{L}_{AB}(t)\psi_0 + \int_0^t \mathfrak{L}_{AB}(t-s)[f(\cdot, \psi(s))]ds.$$

$$\psi(t) = \mathfrak{L}_{AB}(t)\psi_0 + \int_0^t \mathfrak{L}_{AB}(t-s)[f(\cdot, \psi(s))]ds.$$

Assuming existence of a discrete operator (“scheme”) $\mathfrak{L}_{AB}^{disc}(k)$,
time step $k > 0$, that approximates $\mathfrak{L}_{AB}(k)$:

$\psi(t) = \mathfrak{L}_{AB}\psi_0$ solution to the homogeneous equation. Fix $T > 0$.
Then there exist a constant $C > 0$ and an integer $j \geq 1$, such that

$$\|\mathfrak{L}_{AB}^{disc}(k)[\psi(t)] - \psi(t+k)\| \leq Ck^{j+1}, \quad 0 \leq t \leq T.$$

DISCRETIZATION OF NONHOMOGENEOUS EQUATION

Splitting with two “generators”, $A + B$ and f .

$$(i) \quad \psi_t = A[\psi] + B[\psi],$$

$$(ii) \quad \psi_t = f(\cdot, \psi).$$

$$\psi_{tt} = f'_\psi(\cdot, \psi(t)) \cdot \psi_t = f'_\psi(\cdot, \psi(t)) \cdot f(\cdot, \psi(t)).$$

Discretization of (ii):

$$\mathfrak{M}^{disc}(k)[\psi(t)] = \psi(t) + kf(\cdot, \psi(t)) + \frac{k^2}{2} f'_\psi(\cdot, \psi(t)) \cdot f(\cdot, \psi(t)).$$

SUMMARY: A discrete operator $\Gamma^{disc}(k)$ for the approximation of the full system over the time interval $[t, t + k]$ is given by

$$\Gamma^{disc}(k) = \mathfrak{M}^{disc}(k) \mathfrak{L}_{AB}^{disc}(k).$$

A SCALAR MODEL ON MANIFOLDS

- ▶ Good definition of **NONLINEAR VECTORFIELDS** is needed for $u_t + \operatorname{div}F(u) = 0$.

A SCALAR MODEL ON MANIFOLDS

- ▶ Good definition of **NONLINEAR VECTORFIELDS** is needed for $u_t + \operatorname{div}F(u) = 0$.
- ▶ Lack of linear structure (translation invariance)—more difficult to control **TOTAL VARIATION** which is related to L^1 contraction between two translated solutions.

A SCALAR MODEL ON MANIFOLDS

- ▶ Good definition of **NONLINEAR VECTORFIELDS** is needed for $u_t + \operatorname{div}F(u) = 0$.
- ▶ Lack of linear structure (translation invariance)—more difficult to control **TOTAL VARIATION** which is related to L^1 contraction between two translated solutions.
- ▶ No **SELF-SIMILAR SOLUTIONS**—Riemann Problems are not defined.

A SCALAR MODEL ON MANIFOLDS

- ▶ Good definition of **NONLINEAR VECTORFIELDS** is needed for $u_t + \operatorname{div}F(u) = 0$.
- ▶ Lack of linear structure (translation invariance)—more difficult to control **TOTAL VARIATION** which is related to L^1 contraction between two translated solutions.
- ▶ No **SELF-SIMILAR SOLUTIONS**—Riemann Problems are not defined.
- ▶ Waves produce multiple “recurring” interactions.

DEFINITION:

A **flux** on a manifold (M^n, g) is a vector field $f = f_x(\bar{u})$ depending upon the parameter \bar{u} (the dependence in both variables being smooth).

DEFINITION:

A **flux** on a manifold (M^n, g) is a vector field $f = f_x(\bar{u})$ depending upon the parameter \bar{u} (the dependence in both variables being smooth).

The **conservation law associated with the flux** f_x on M is

$$\partial_t u + \nabla_g \cdot (f_x(u)) = 0,$$

Unknown: scalar-valued function $u = u(t, x)$.

$\nabla_g \cdot (f_x(u))$ for fixed t , on vector field $x \mapsto f_x(u(t, x)) \in T_x M$.

DEFINITION:

A **flux** on a manifold (M^n, g) is a vector field $f = f_x(\bar{u})$ depending upon the parameter \bar{u} (the dependence in both variables being smooth).

The **conservation law associated with the flux** f_x on M is

$$\partial_t u + \nabla_g \cdot (f_x(u)) = 0,$$

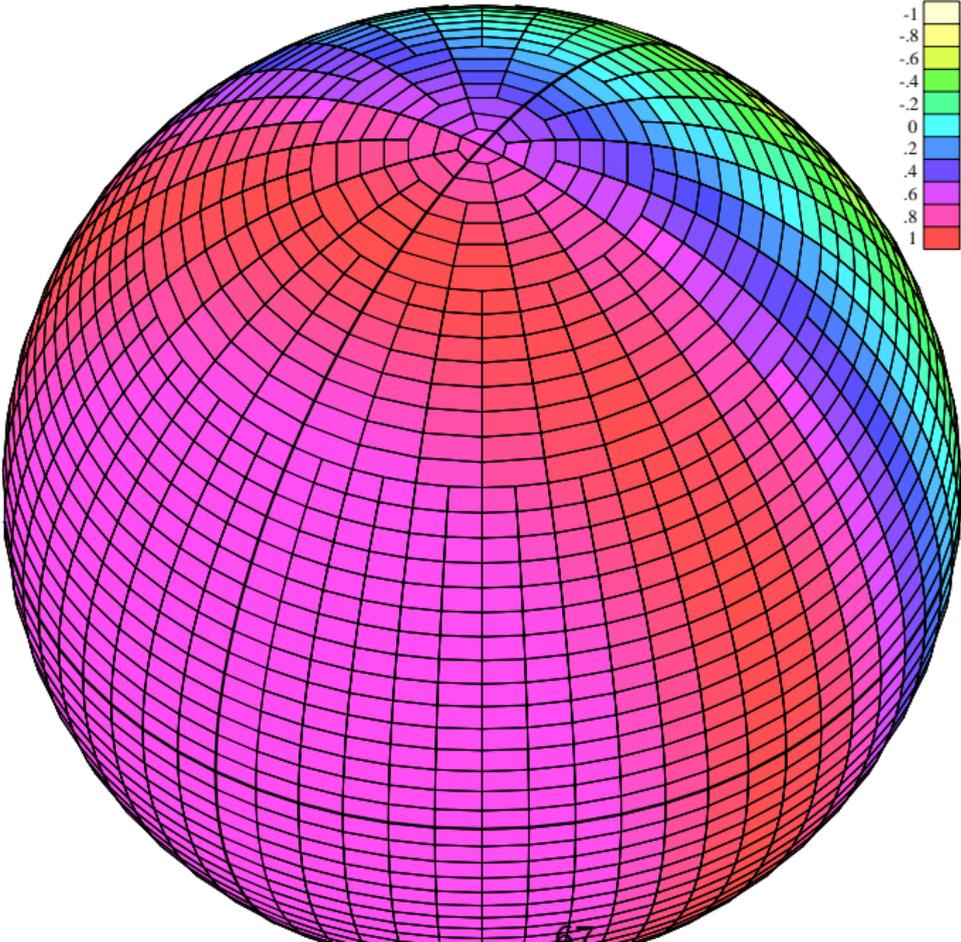
Unknown: scalar-valued function $u = u(t, x)$.

$\nabla_g \cdot (f_x(u))$ for fixed t , on vector field $x \mapsto f_x(u(t, x)) \in T_x M$.

A flux is called **geometry-compatible** if it satisfies the divergence-free condition

$$\nabla \cdot f_x(\bar{u}) = 0, \quad \bar{u} \in \mathbf{R}, x \in M.$$

CONFINED SOLUTION



S P H E R I E . 0 6 // c c j l S a t M a y 1 0 1 9 2 7 : 2 6 2 0 0 8

The regularized initial-value problem

An initial data $u_0 \in BV(M; dV_g)$, find a solution $u^\varepsilon = u^\varepsilon(t, x)$ to:

$$\partial_t u^\varepsilon + \operatorname{div}_g (f_x(u^\varepsilon)) = \varepsilon \Delta_g u^\varepsilon, \quad x \in M, t \geq 0,$$

$$u^\varepsilon(0, x) = u_0^\varepsilon(x), \quad x \in M,$$

where Δ_g denotes the Laplace operator on the manifold M ,

$$\begin{aligned} \Delta_g v &:= \nabla_g \cdot \nabla_g v \\ &= g^{ij} \left(\frac{\partial^2 v}{\partial x^i \partial x^j} - \Gamma_{ij}^k \frac{\partial v}{\partial x^k} \right). \end{aligned}$$

$u_0^\varepsilon : M \rightarrow \mathbf{R}$ is a sequence of smooth functions satisfying

$$\|u_0^\varepsilon\|_{L^p(M)} \leq \|u_0\|_{L^p(M)}, \quad p \in [1, \infty],$$

$$TV(u_0^\varepsilon) \leq TV(u_0),$$

$$\sup_{0 < \varepsilon < 1} \varepsilon \|u_0^\varepsilon\|_{H^2(M; dV_g)} < \infty,$$

$$u_0^\varepsilon \rightarrow u_0 \quad \text{a.e. on } M.$$

REGULARIZED PROBLEM

(Ben-Artzi and LeFloch, 2007)

THEOREM: Let $f = f_x(\bar{u})$ be a geometry-compatible flux on (M, g) . Given any initial data $u_0^\varepsilon \in C^\infty(M)$ satisfying the above conditions there exists a unique solution $u^\varepsilon \in C^\infty(\mathbf{R}_+ \times M)$ to the initial value problem. Moreover, for each $1 \leq p \leq \infty$ the solution satisfies

$$\|u^\varepsilon(t)\|_{L^p(M; dV_g)} \leq \|u^\varepsilon(t')\|_{L^p(M; dV_g)}, \quad 0 \leq t' \leq t$$

and, for any two solutions u^ε and v^ε ,

$$\begin{aligned} & \|v^\varepsilon(t) - u^\varepsilon(t)\|_{L^1(M; dV_g)} \\ & \leq \|v^\varepsilon(t') - u^\varepsilon(t')\|_{L^1(M; dV_g)}, \quad 0 \leq t' \leq t. \end{aligned}$$

In addition, for every convex entropy/entropy flux pair (U, F_x) the solution u^ε satisfies the entropy inequality

$$\partial_t U(u^\varepsilon) + \operatorname{div}_g (F_x(u^\varepsilon)) \leq \varepsilon \Delta_g U(u^\varepsilon).$$

ENTROPY SOLUTION

(Ben-Artzi and LeFloch, 2007)

CORRECTION: Lengeler and Müller 2013

THEOREM: Let $f = f_x(\bar{u})$ be a geometry-compatible flux on (M, g) . Given any bounded initial function $u_0 \in BV(\mathbf{M}^n; dV_g)$ there exists an **entropy solution** $u \in L^\infty(\mathbf{R}_+ \times \mathbf{M}^n)$ to the initial value problem, so that

$$\|u(t)\|_{L^p(\mathbf{M}^n; dV_g)} \leq \|u_0\|_{L^p(\mathbf{M}^n; dV_g)}, \quad t \geq 0, p \in [1, \infty].$$

For some constant $C_1 > 0$ depending on $\|u_0\|_{L^\infty(M)}$ and the Ricci tensor

$$\begin{aligned} TV(u(t)) &\leq e^{C_1 t} (1 + TV(u_0)), \quad t \in \mathbf{R}_+, \\ \|u(t) - u(t')\|_{L^1(M; dV_g)} &\leq C_1 TV(u_0) |t - t'|, \quad 0 \leq t' \leq t. \end{aligned} \quad (1)$$

Definition

Let $f = f_x(\bar{u})$ be a geometry-compatible flux on (M, g) . Given any initial condition $u_0 \in L^\infty(M)$, a measure-valued map $(t, x) \in \mathbf{R}_+ \times M \mapsto \nu_{t,x}$ is called an **entropy measure-valued solution** to the initial value problem if, for every convex entropy/entropy flux pair (U, F_x) ,

$$\begin{aligned} & \iint_{\mathbf{R}_+ \times M} \left(\langle \nu_{t,x}, U \rangle \partial_t \theta(t, x) + \right. \\ & \left. g_x(\langle \nu_{t,x}, F_x \rangle, \text{grad}_g \theta(t, x)) \right) dV_g(x) dt \\ & + \int_M U(u_0(x)) \theta(0, x) dV_g(x) \geq 0, \end{aligned} \quad (2)$$

for every smooth function $\theta = \theta(t, x) \geq 0$ compactly supported in $[0, +\infty) \times M$.

THEOREM

(Well-posedness theory in the measure-valued class for geometry-compatible conservation laws.)

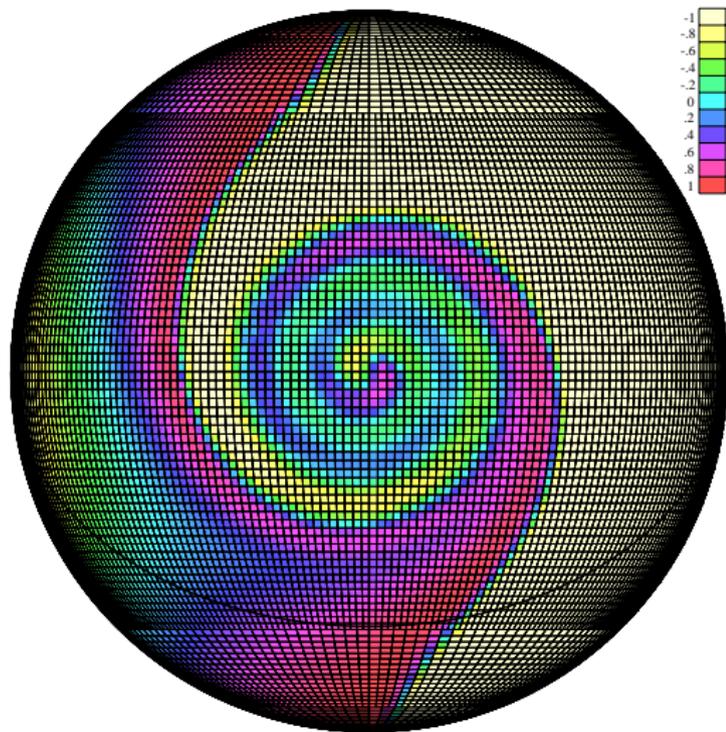
(Ben-Artzi and LeFloch 2006)

Let $f = f_x(\bar{u})$ be a geometry-compatible flux on (M, g) , and let $u_0 \in L^\infty(M)$. Then there exists a unique entropy measure-valued solution ν to the initial value problem. For almost every (t, x) , the measure $\nu_{t,x}$ is a Dirac mass, i.e. of the form

$$\nu_{t,x} = \delta_{u(t,x)},$$

where the function $u \in L^\infty(\mathbf{R}_+ \times M)$. Moreover, the initial data is attained in the strong sense

$$\limsup_{t \rightarrow 0^+} \int_M |u(t, x) - u_0(x)| dV_g(x) = 0. \quad (3)$$



SPHERE OF f (cf. MonApr 13.000013.2016)

THANK YOU!